
Spectroscopie haute résolution et mesures de fréquences

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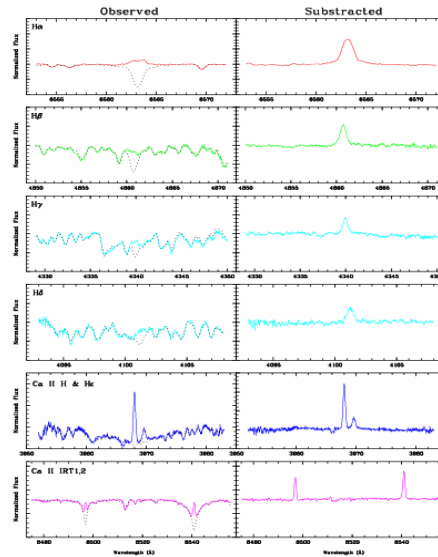
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Exemples

Spectroscopic Analysis

- **Kinematics (U, V, W).**
 - Radial velocity (V_r)
- **Age (LiI 6707.8Å).**
- **Chromospheric activity**
 - CaII H&K to CaII IRT
- **Rotation (v_{seni}).**
 - Activity – rotación relation
- **Stellar parameters.**
 - T_{eff} , $\log g$, ξ and $[\text{Fe}/\text{H}]$
- **Absolute and differential abundances.**
 - Chemical tagging



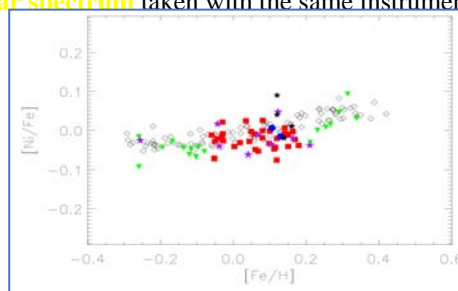
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Chemical abundances

Fe, Na, Mg

Fe, Na, Mg, Al, Si, Ca, Sc, Ti, V, Cr, Mn, Co, and Ni

- **EW method** in a line-by-line basis with *ARES* code (Sousa et al. 2007).
- **Line lists and atomic parameters** from (Neves et al. 2009; González Hernández et al. 2010).
- Abundance analysis with *MOOG* (Snedden 1973) using our determined atmospheric parameters and a **solar spectrum** taken with the same instrumental configuration.



$[\text{Ni}/\text{Fe}]$ vs $[\text{Fe}/\text{H}]$: open diamonds represent the thin disk data (González Hernández et al. 2010), black filled triangles represent Hyades cluster data (Paulson et al. 2003). **Red points** are our stars compatible with Hyades Fe abundance, and the **green** ones not compatible. BZ Cet and HD19902 Hyades cluster members are marked with **blue circles**. Purple starred points represent the giant stars. Black starred points are the candidates selected stars in De Silva et al. (2011), black circles are those selected in Pompéia et al. (2011).

Plan

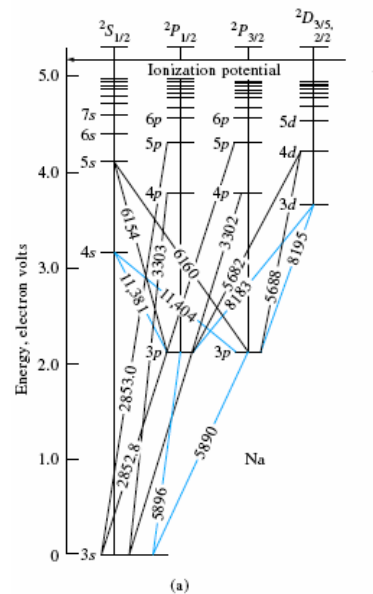
1. Spectroscopie haute résolution
2. Effets d'élargissements
3. Méthodes « Doppler-free »
 - 2 photon
 - Absorption saturée
4. Mesure de fréquence
 - spectre de I_2
 - monochromateur
 - détection optogalvanique
 - lambda-mètre,
 - chaîne de fréquence
 - peigne de fréquence

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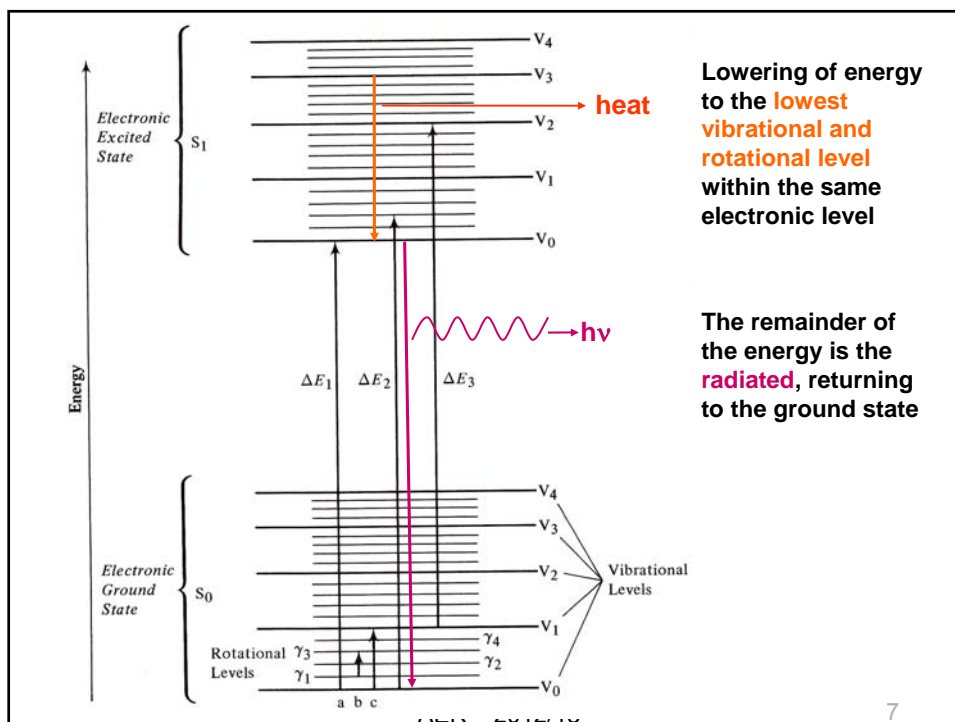
1. Spectroscopie haute résolution

- **Spectroscopie atomique:** peu de raies, typiquement énergie élevées (visible, UV)
- bcp de données connues, aussi de l'astrophysique
- **Spectroscopie moléculaire:** mélange de transitions électronique et ro-vibrationnelle
- situation bcp plus complexe



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1. Spectroscopie haute résolution

→ Å, nm, cm^{-1} , (kayser), MHz

→ haute résolution temporelle – spectrale

$$\Delta\nu = (2\pi\tau)^{-1}$$

$$\Delta\nu = \frac{1}{2\pi\tau}$$

largeur de raie naturelle

→ principe d'incertitude de Heisenberg pour un paquet d'onde

$$\Delta x \cdot \Delta p_x \geq \hbar/2 \quad \text{ou} \quad \leftrightarrow \quad \Delta E \cdot \Delta t \geq \hbar$$

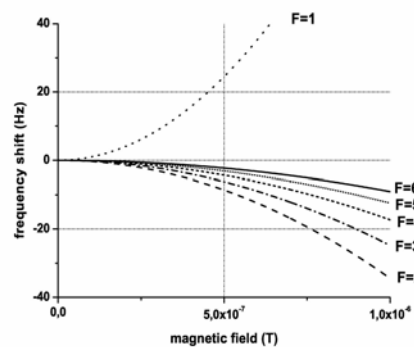
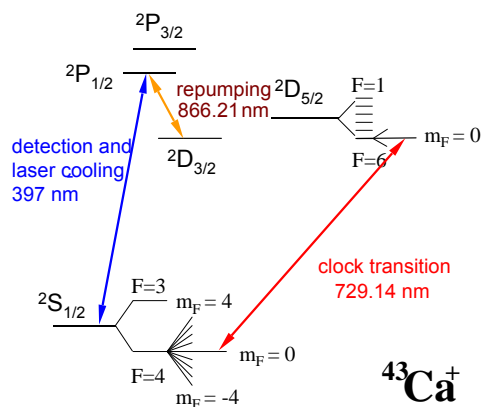
2. Effets d'élargissement

- (1) Dans un jet : temps de transit
- (2) Agitation thermique : effet Doppler
- (3) Modification de la durée de vie : collisions
- (4) Par un champ externe : Zeeman, Stark
- (5) Par saturation
- (6) ...

$$\frac{\Delta\nu}{\nu} = 7.16 \times 10^{-7} \sqrt{\frac{T}{M}}$$

Effets d'élargissement et de déplacement (broadening and shift)

Line broadening and shift



Effets systématiques - $^{43}\text{Ca}^+$

effet	proportionnel à	condition	δf [Hz]	$\delta f/f$
DC Stark* (corps noir)	$5.7 + 2.1 \times \frac{1}{2} (3\cos^2\theta - 1)$	1V/mm T=300K	0.39±0.27	9.5(±6.5)x10⁻¹⁶
quadrupole shift	$-8.1 \times 10^{-7} \left(\frac{1}{2} \frac{\partial^2 V}{\partial x^2} \right) \Pi$	1V/mm sur 1mm (piège sphérique) 3 mesures	+/-0.1	+/- 2.5x10⁻¹⁶
AC Stark (dépl.lumineux)		0.75μW/mm² @ 729nm et 0.1μT	±0.006	± 1.5x10⁻¹⁷
Doppler 2 nd ordre	$- \frac{1}{2} v^2/c^2$	Doppler limit, <n>≈10	-1x10⁻⁴	-2.5x10⁻¹⁹
Zeeman linéaire			0	0
Zeeman quadr.*	$-9.05 \text{ Hz}/\mu\text{T}^2$	(0.1±0.05 μT) stability	0.09±0.09	-2.2(±2.2)x10⁻¹⁶
TOTAL			+0.3±0.4	(+7±10)x10⁻¹⁶

* utilisation de $^2\text{S}_{1/2}$, F=4, m_F=0 → $^2\text{D}_{5/2}$, F=6, m_F=0

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Broadening and shifts - experimental constraints -

$$\Delta\nu_{\text{nat}}/\nu \approx 3 \cdot 10^{-16}$$

to reach a 1-Hz linewidth at 411 THz ($\Delta\nu/\nu = 2.4 \cdot 10^{-15}$)

ion at the Doppler limit, <n>≈10

2nd order Doppler

magnetic field < (0.1 ± 0.05 μT)

quadratic Zeeman, AC Stark

residual electric field <1V/mm

DC Stark

electric field gradient <1V/mm over 1 mm
(spherical trap)

quadrupole moment

3 optical axes (orthogonal)

measure of the quadrupole
moment

T= 300K

DC Stark (black-body radiation)

P<0.75μW/mm² @
729nm

AC Stark (light shift)

B



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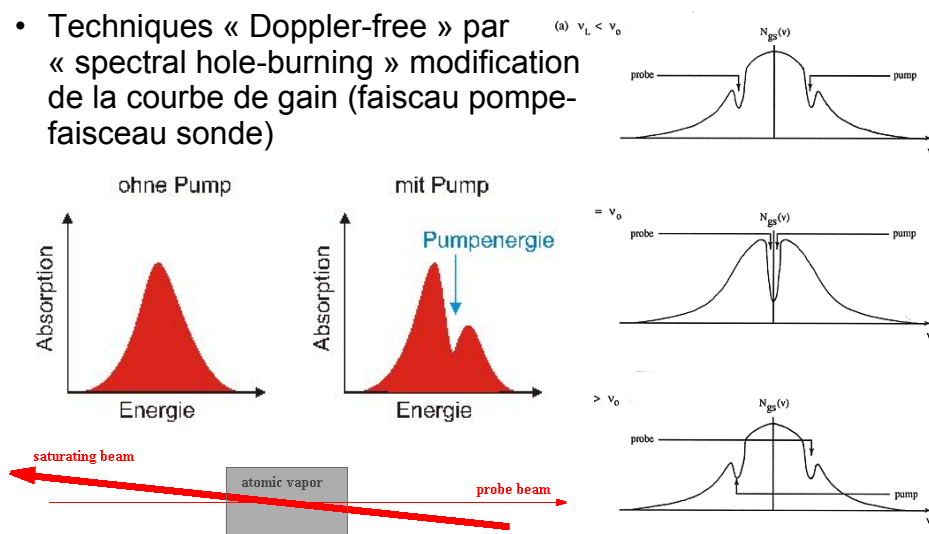
3. Méthodes sans effet Doppler

Techniques non-linéaires

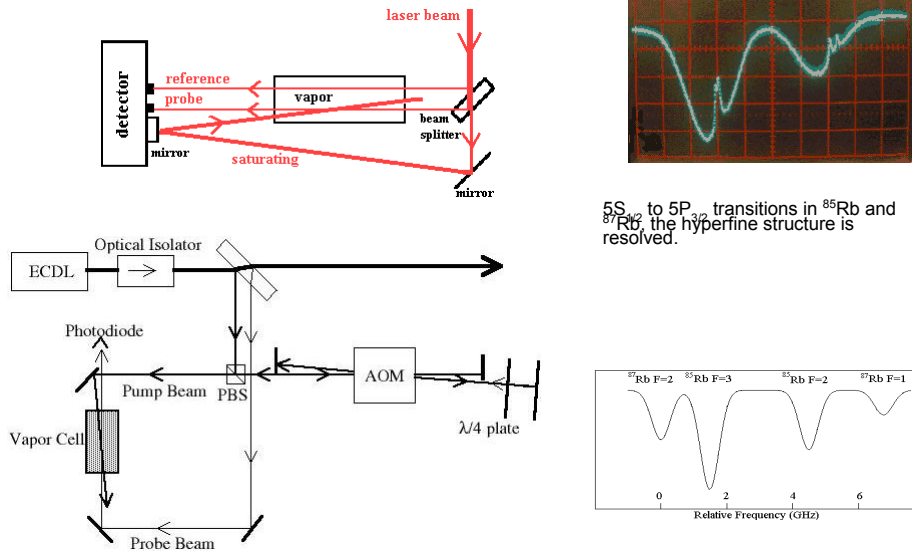
- Absorption saturée
- Spectroscopie à 2 photons
- ...

Absorption saturée

- Techniques « Doppler-free » par « spectral hole-burning » modification de la courbe de gain (faisceau pompe-faisceau sonde)



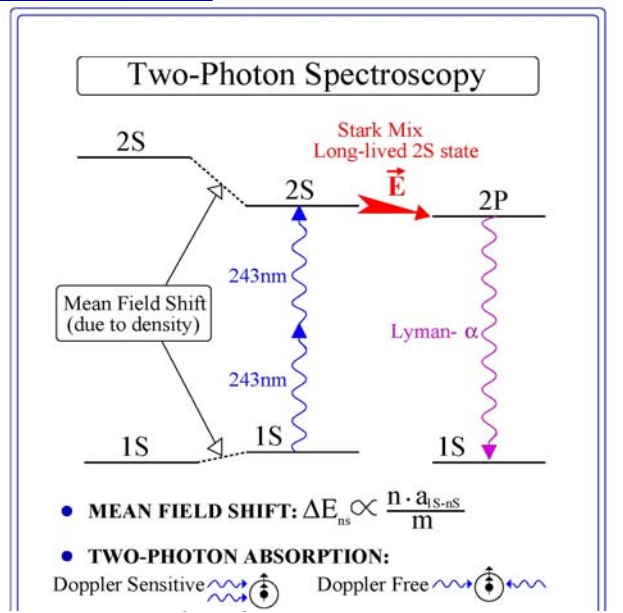
Absorption saturée



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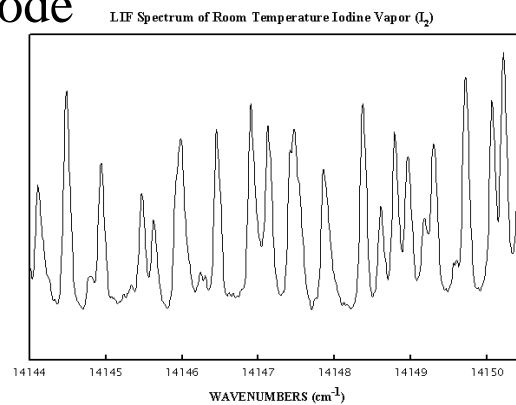
2 photons spectroscopie



Mesures de fréquence

Mesure de fréquence

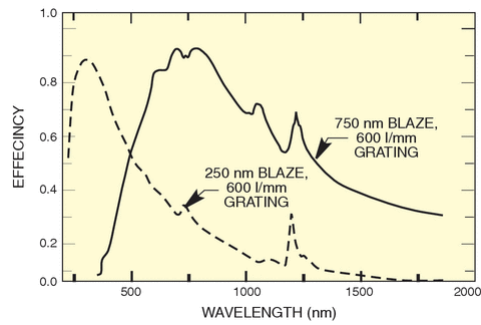
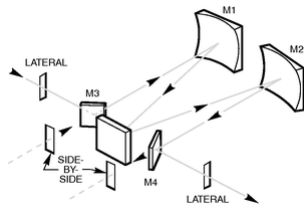
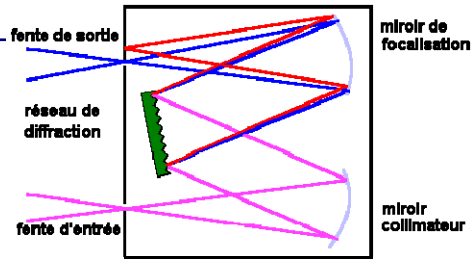
- spectre de l'iode



Mesure de fréquence

• Monochromateur

resolution 1nm

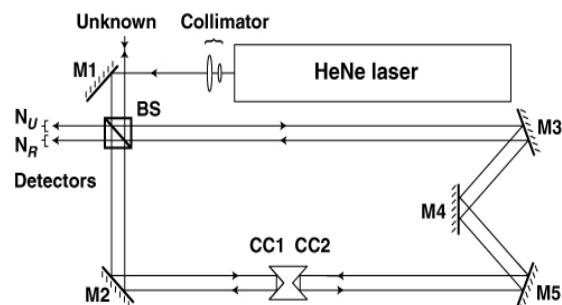


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Mesure de fréquence - λ -mètre

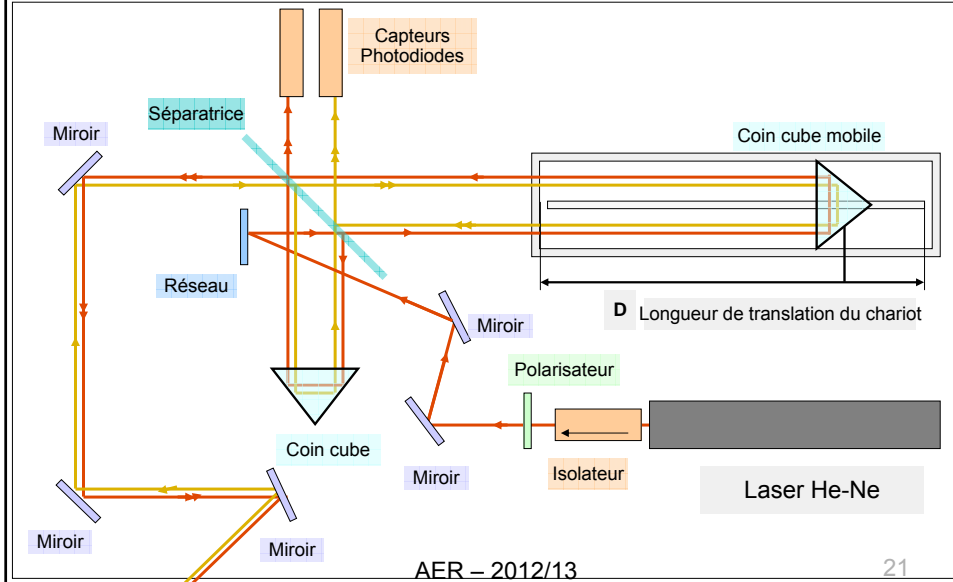
- Ratio inverse du nombre des franges d'interférences
- Précision 10^{-5} à 10^{-8}



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Schéma d'ensemble du Lambda mètre.



Mesure de fréquence – chaîne de fréquence

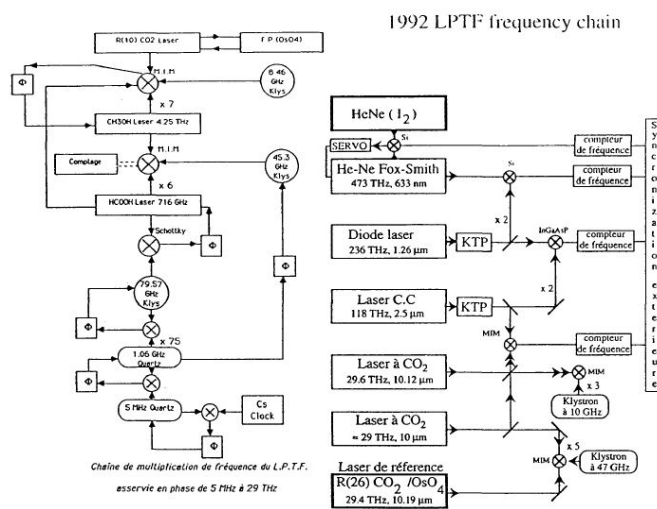
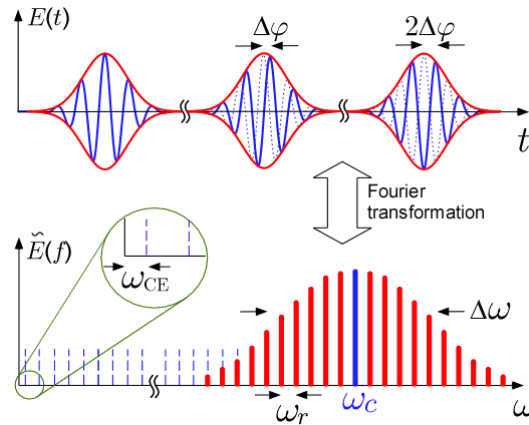


Figure 2 : chaîne de fréquences reliant l'horloge à césium vers 9 GHz à la région spectrale de 10 μm [2].

Figure 3 : chaîne de fréquences reliant la région spectrale de 10 μm à la fréquence visible [4]

Peigne de fréquence

Laser TiSa fs
+ fibre non-linéaire



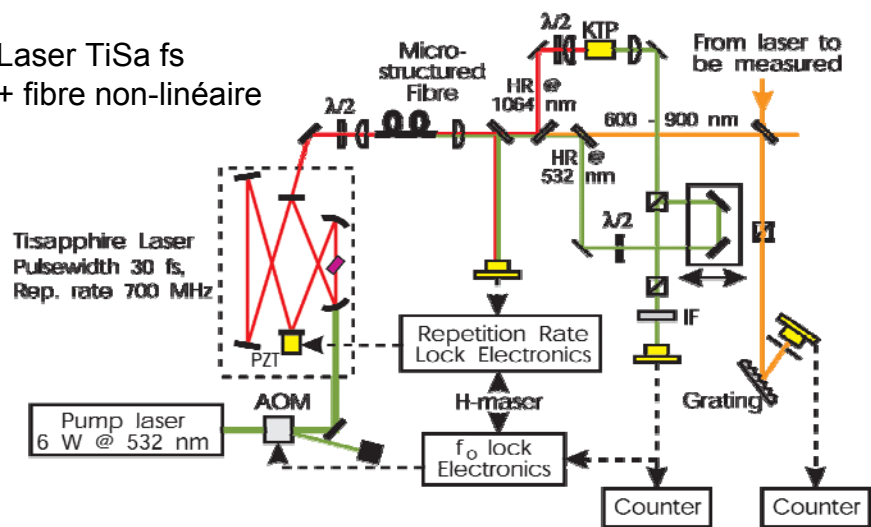
Three consecutive pulses of the pulse train emitted by a mode locked laser and the corresponding spectrum. The pulse-to-pulse phase shift results in a offset frequency $\text{CE} = \omega_c - \omega_r$ because the optical carrier wave at C moves with the phase velocity while the envelope moves with the group velocity.

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Peigne de fréquence

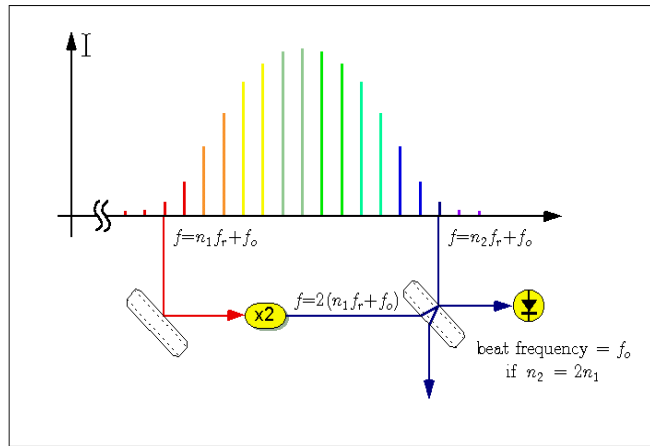
Laser TiSa fs
+ fibre non-linéaire



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Peigne de fréquence - *autoreféréncement*



Th.Udem, R.Holzwarth, and T.W.Hänsch, Nature 416, 233 (2002)

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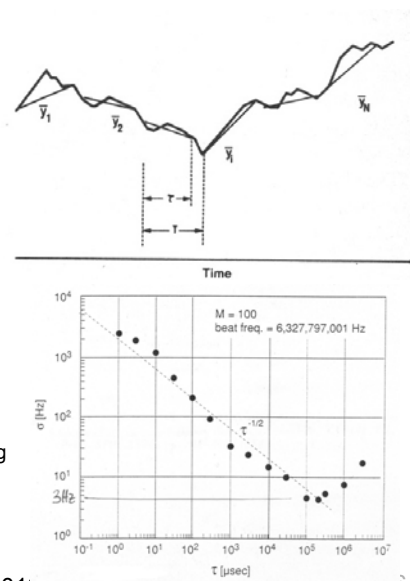
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Time domain (Allen variance)

Integration of measurement data over different times τ

$$\sigma_y^2(\tau) = \frac{1}{2} \langle (\bar{y}_2 - \bar{y}_1)^2 \rangle$$

Here: maximal stability of 3 Hz at integrating time of 0.2 sec.



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